Spatio-temporal comparison of initial perturbation methods

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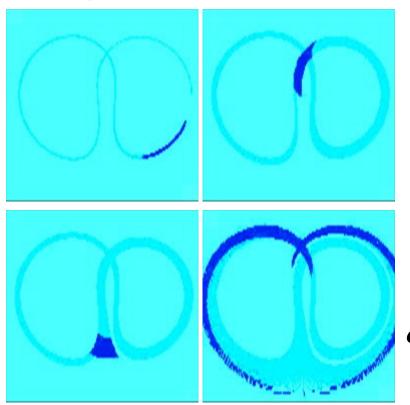


Outline

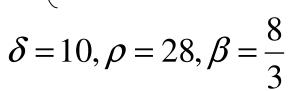
- Uncertainty in the initial conditions
- Perturbations growth in toy models
- Interface theory and MVL diagram
- Comparison of initial perturbations techniques in NWP models
- Summary and conclusions

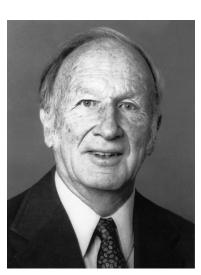
Chaos in the atmosphere: Lorenz model

Chaos was introduced by Edward Lorenz in 1963, when he used a simplified system to model the earth's weather on a computer.



$$\begin{cases} \frac{dx}{dt} = \delta(y - x) \\ \frac{dy}{dt} = \rho x - y - xz \\ \frac{dz}{dt} = xy - \beta z \end{cases}$$





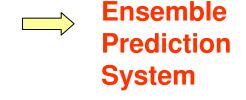
"... one flap of a sea-gull's wing may forever change the future course of the weather" (Lorenz, 1963)

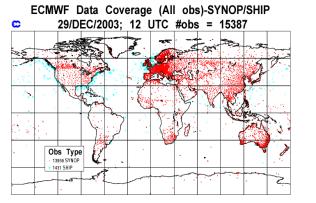
Small errors in the initial conditions become larger in time.

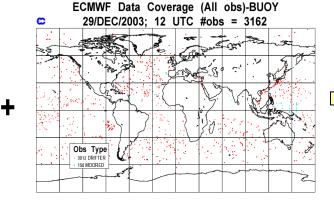
Sources of Uncertainty:

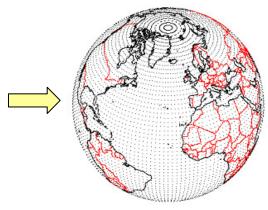
Uncertainty in the initial conditions:

- sampling error in measurements
- systematic error in measurements (e.g. instrumental biases)
- etc.









Uncertainty in the model:

- Physical uncertainty in model parameters
- Sampling uncertainty in statistical estimates of parameters



- Incomplete knowledge of external factors
- etc.

Toy models: Non linear dynamics & fractal geometry

Simple mathematical models allow us to understand the complex dynamic of nonlinear systems.

Discrete case:

 Iterative
 systems

 Logistic maps

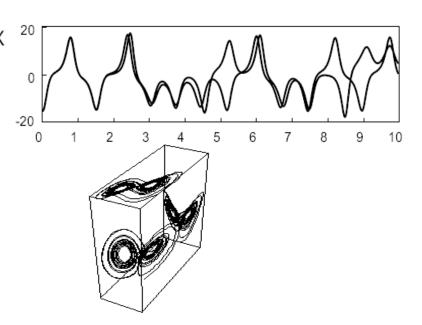
$$x_{n+1} = 4x_n (1-x_n)_{0.6}^{0.8} \left[\begin{array}{c} 0.8 \\ 0.4 \\ 0.2 \\ 0 \end{array} \right]_{0.2000} \left[\begin{array}{c} 0.8 \\ 0.4 \\ 0.2 \\ 0 \end{array} \right]_{0.2000} \left[\begin{array}{c} 0.8 \\ 0.4 \\ 0.2 \\ 0 \end{array} \right]_{0.2000} \left[\begin{array}{c} 0.8 \\ 0.4 \\ 0.2 \\ 0 \end{array} \right]_{0.2000} \left[\begin{array}{c} 0.8 \\ 0.4 \\ 0.2 \\ 0 \end{array} \right]_{0.2000} \left[\begin{array}{c} 0.8 \\ 0.4 \\ 0.2 \\ 0 \end{array} \right]_{0.2000} \left[\begin{array}{c} 0.8 \\ 0.4 \\ 0.2 \\ 0 \end{array} \right]_{0.2000} \left[\begin{array}{c} 0.8 \\ 0.4 \\ 0.2 \\ 0 \end{array} \right]_{0.2000} \left[\begin{array}{c} 0.8 \\ 0.4 \\ 0.2 \\ 0 \end{array} \right]_{0.2000} \left[\begin{array}{c} 0.8 \\ 0.4 \\ 0.2 \\ 0 \end{array} \right]_{0.2000} \left[\begin{array}{c} 0.8 \\ 0.4 \\ 0.2 \\ 0 \end{array} \right]_{0.2000} \left[\begin{array}{c} 0.8 \\ 0.4 \\ 0.2 \\ 0 \end{array} \right]_{0.2000} \left[\begin{array}{c} 0.8 \\ 0.4 \\ 0.2 \\ 0 \end{array} \right]_{0.2000} \left[\begin{array}{c} 0.8 \\ 0.4 \\ 0.2 \\ 0 \end{array} \right]_{0.2000} \left[\begin{array}{c} 0.8 \\ 0.4 \\ 0.2 \\ 0 \end{array} \right]_{0.2000} \left[\begin{array}{c} 0.8 \\ 0.4 \\ 0.2 \\ 0 \end{array} \right]_{0.2000} \left[\begin{array}{c} 0.8 \\ 0.4 \\ 0.2 \\ 0 \end{array} \right]_{0.2000} \left[\begin{array}{c} 0.8 \\ 0.4 \\ 0.2 \\ 0 \end{array} \right]_{0.2000} \left[\begin{array}{c} 0.8 \\ 0.4 \\ 0.2 \\ 0 \end{array} \right]_{0.2000} \left[\begin{array}{c} 0.8 \\ 0.4 \\ 0.2 \\ 0 \end{array} \right]_{0.2000} \left[\begin{array}{c} 0.8 \\ 0.4 \\ 0.2 \\ 0 \end{array} \right]_{0.2000} \left[\begin{array}{c} 0.8 \\ 0.4 \\ 0.2 \\ 0 \end{array} \right]_{0.2000} \left[\begin{array}{c} 0.8 \\ 0.4 \\ 0.2 \\ 0 \end{array} \right]_{0.2000} \left[\begin{array}{c} 0.8 \\ 0.4 \\ 0.2 \\ 0 \end{array} \right]_{0.2000} \left[\begin{array}{c} 0.8 \\ 0.4 \\ 0.2 \\ 0 \end{array} \right]_{0.2000} \left[\begin{array}{c} 0.8 \\ 0.4 \\ 0.2 \\ 0 \end{array} \right]_{0.2000} \left[\begin{array}{c} 0.8 \\ 0.4 \\ 0.2 \\ 0 \end{array} \right]_{0.2000} \left[\begin{array}{c} 0.8 \\ 0.4 \\ 0.2 \\ 0 \end{array} \right]_{0.2000} \left[\begin{array}{c} 0.8 \\ 0.4 \\ 0.2 \\ 0 \end{array} \right]_{0.2000} \left[\begin{array}{c} 0.8 \\ 0.4 \\ 0.2 \\ 0 \end{array} \right]_{0.2000} \left[\begin{array}{c} 0.8 \\ 0.4 \\ 0.2 \\ 0.2 \end{array} \right]_{0.2000} \left[\begin{array}{c} 0.8 \\ 0.4 \\ 0.2 \\ 0.2 \end{array} \right]_{0.2000} \left[\begin{array}{c} 0.8 \\ 0.4 \\ 0.2 \\ 0.2 \end{array} \right]_{0.2000} \left[\begin{array}{c} 0.8 \\ 0.4 \\ 0.2 \\ 0.2 \end{array} \right]_{0.2000} \left[\begin{array}{c} 0.8 \\ 0.4 \\ 0.2 \\ 0.2 \end{array} \right]_{0.2000} \left[\begin{array}{c} 0.8 \\ 0.4 \\ 0.2 \\ 0.2 \end{array} \right]_{0.2000} \left[\begin{array}{c} 0.8 \\ 0.4 \\ 0.2 \\ 0.2 \end{array} \right]_{0.2000} \left[\begin{array}{c} 0.8 \\ 0.4 \\ 0.2 \\ 0.2 \end{array} \right]_{0.2000} \left[\begin{array}{c} 0.8 \\ 0.4 \\ 0.2 \\ 0.2 \end{array} \right]_{0.2000} \left[\begin{array}{c} 0.8 \\ 0.4 \\ 0.2 \\ 0.2 \end{array} \right]_{0.2000} \left[\begin{array}{c} 0.8 \\ 0.4 \\ 0.2 \\ 0.2 \end{array} \right]_{0.2000} \left[\begin{array}{c} 0.8 \\ 0.4 \\ 0.2 \\ 0.2 \end{array} \right]_{0.2000} \left[\begin{array}{c} 0.8 \\ 0.4 \\ 0.2 \\ 0.2 \end{array} \right]_{0.2000} \left[\begin{array}{c} 0.8 \\ 0.4 \\ 0.2 \\ 0.2 \end{array} \right]_{0.2000} \left[\begin{array}{c} 0.8 \\ 0.4 \\ 0.2 \end{array} \right]_{0.2000} \left[\begin{array}{c} 0.8 \\ 0.4 \\ 0.2 \end{array} \right]_{0.2000} \left[\begin{array}{c} 0.8 \\ 0.4 \\ 0.2 \end{array} \right]_{0.2000} \left[\begin{array}{c} 0.8 \\ 0.4 \\$$

Continuous case:
 Differential
 equations

$$\frac{dx}{dt} = \delta(y - x)$$

$$\frac{dy}{dt} = \rho x - y - xz$$

$$\frac{dz}{dt} = xy - \beta z$$



Toy models: Spatial coupling

Spatial points do not grow independently

Coupled chaotic maps:

$$u(x,t+1) = \mathcal{E}f(u(x+1,t)) + \mathcal{E}f(u(x-1,t)) + (1-2\mathcal{E})f(u(x,t))$$

$$f(u) = 4u(1-u)$$

Logistic map

$$u(x+L,t) = u(x,t), x = 1,...,L$$
 Boundary conditions

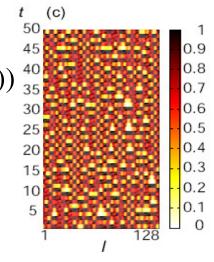
$$\varepsilon = \frac{1}{3}$$

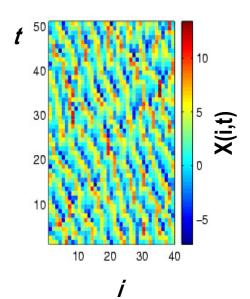
Homogeneous coupling

Lorenz-96 model:

$$\frac{dx_i}{dt} = -x_{i-1}(x_{i-2} - x_{i+1}) - x_i + F, \quad i = 1, ..., L$$

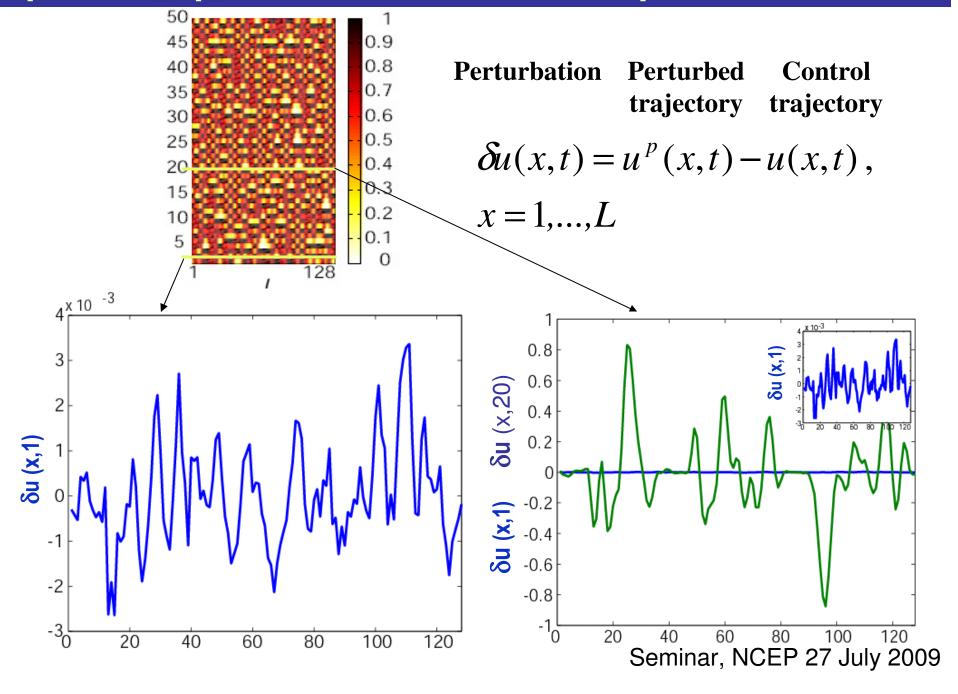
This model simulates the evolution of an atmospheric variable x in L grid points over a latitude





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Spatio-temporal characterization of perturbations



Temporal growth of the perturbations

 Perturbations will exhibit a long-term average exponential growth with a characteristic rate:

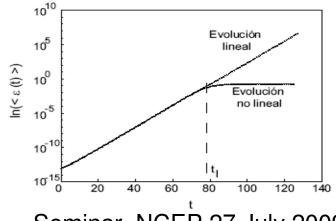
$$| \delta u(x,t) | \approx e^{\lambda_1 t}$$
 where λ_1 is the leading Lyapunov exponent

$$\mathcal{E}(t) \equiv \prod_{x=1}^{L} |\delta u(x,t)|^{\frac{1}{L}} \int_{0.05}^{\infty} \int_{0.05}^{0.15} \int_{0.05}^{0.05} \int_{0$$

The lower the initial amplitude of the perturbations, the later they saturate.

In non-linear systems, perturbations become as large as the amplitude of the system.

When the system is linearized they do not saturate.



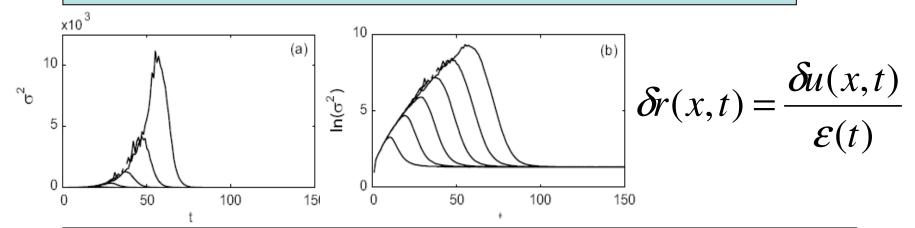
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Spatial growth of the perturbations

$$\sigma_{\delta u}^{2} = \frac{1}{L} \sum_{x=1}^{L} (\delta u(x,t) - \overline{\delta u(x,t)})^{2}$$

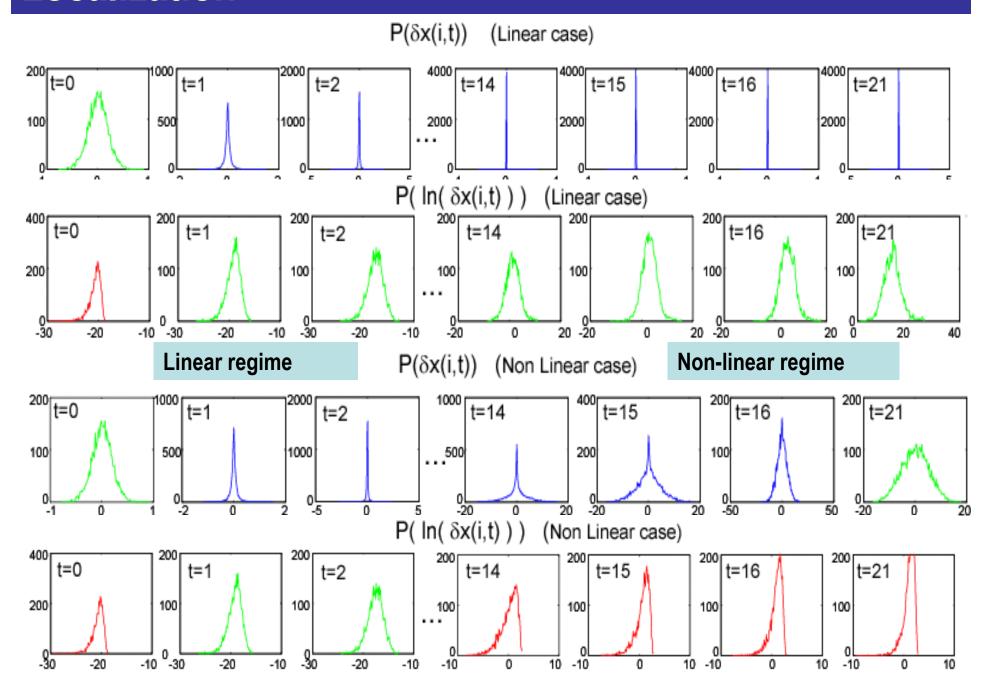
$$\int_{0.05}^{0.1} \int_{0.05}^{0.1} \int_{0.0}^{0.15} \int_{0.0}^{0.15} \int_{0.05}^{0.15} \int_{0.05}^{0.05} \int_{0.05}^{0.05}$$

Dominated by the exponential growth. To avoid it:



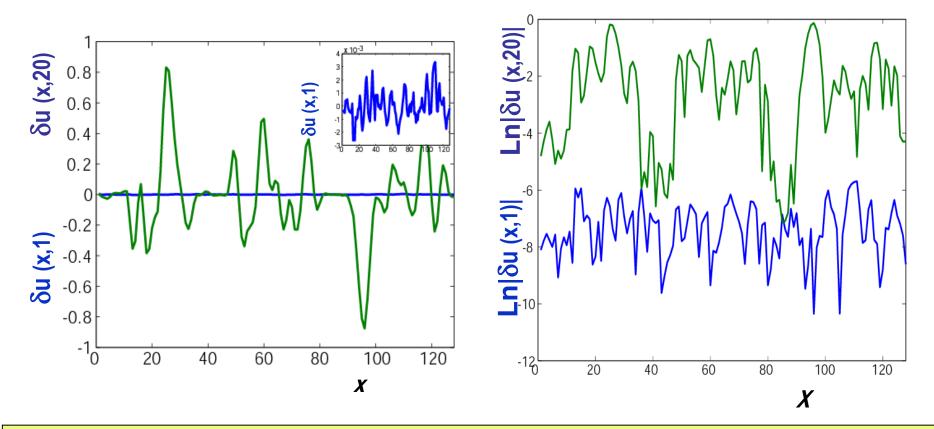
In the non linear regime, perturbations loose variance

Localization



Logarithmic transformation

 $h(x,t) = \ln |\delta u(x,t)|, x = 1,...,L$ Hopf - Cole transformation



There is a strong parallel between the logarithmic perturbations and the development of a rough interfase between two media.

Link between logarithmic perturbations and rough interfaces

- Pikovsky & Kurths (1994) use this transformation to introduce roughening interfaces in the dynamics of perturbations of spatiotemporal chaos.
- Pikovsky & Politi (1998) apply it to study the dynamic localization of Lyapunov vectors in spacetime chaos.
- López et al. (2004) apply this correspondence to explain the scaling properties of growing noninfinitesimal perturbations in space-time chaos.
- Primo et al. (2006) introduce this link to characterize the dynamic scaling of bred vectors in spatially extended chaotic systems.
- Primo et al. (2007) show that the correspondence is also valid for the complex atmospheric circulation models used in weather forecasting.

Interface growth

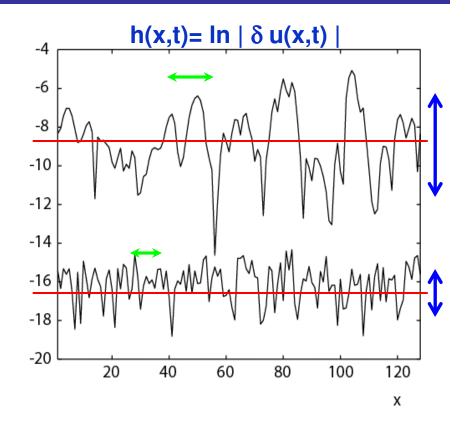
Mean height:
$$\overline{h}(t) = \frac{1}{L} \sum_{x=1}^{L} h(x,t)$$

Width (standard deviation of the height):

$$w(t) = \sqrt{\frac{1}{L} \sum_{x=1}^{L} (h(x,t) - \overline{h}(t))^{2}}$$

Correlation:

$$G(t,l) = \frac{1}{L} \sum_{x=1}^{L} (h(x+l,t) - \overline{h}(t))^{2}$$



Width (Vertical growth)

$$w(t) \approx t^{\beta}$$

$$w(t) \approx l(t)^{\alpha}$$

Correlation length (horizontal growth):

$$l(t) \approx t^{1/z}$$

Barabasi, A.-L. & Stanley, H. E. (1995): "Fractal Concepts in Surface Growth". Cambridge University Press.

Mean-variance of logarithms diagram (MVL)

$$M(t) = \overline{h}(t) = \frac{1}{L} \sum_{x=1}^{L} h(x,t) = \frac{1}{L} \sum_{x=1}^{L} \ln |\delta(x,t)| = \ln \varepsilon(t) \approx \lambda_1 t + \dots$$

$$V(t) = w^2(t) = \frac{1}{L} \sum_{x=1}^{L} (h(x,t) - \overline{h}(t))^2 \approx l(t)^{2\alpha}$$

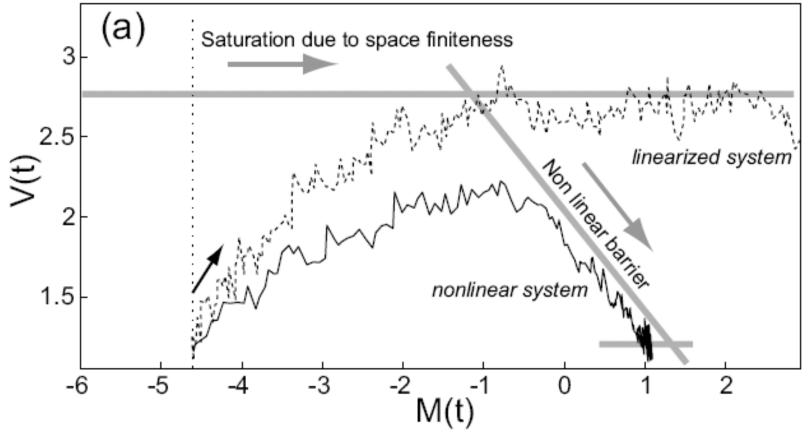


Figure: Gutierrez et al. (2008, Fig 2 (a))

MVL diagram

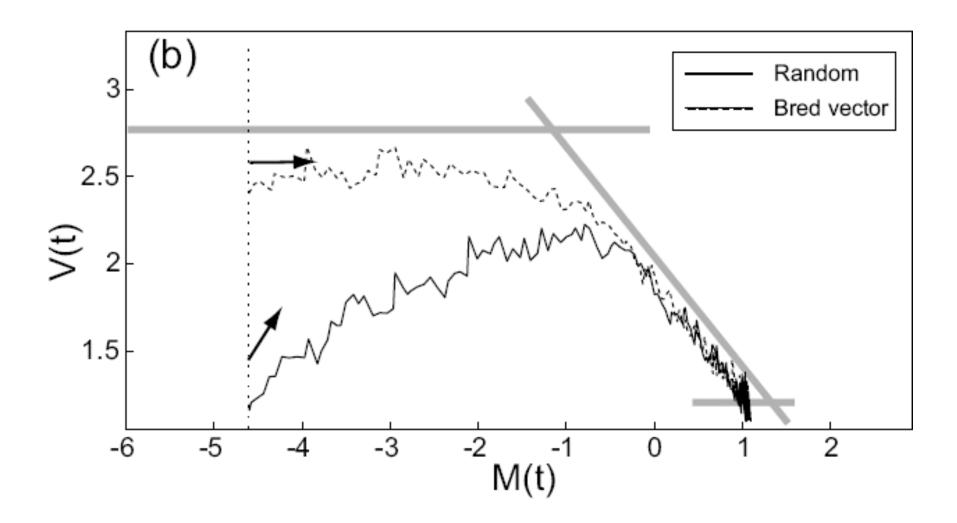
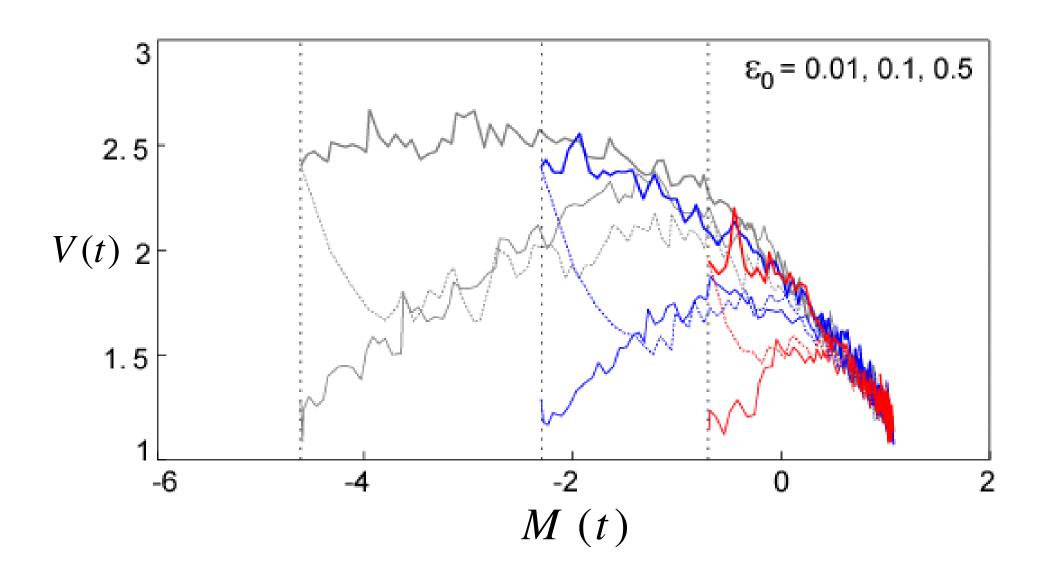


Figure: Gutierrez et al. (2008, Fig 2 (b))

MVL diagram



Numerical models

- Operational forecast Centres represent uncertainty in the initial conditions by perturbing the best estimation of the initial state or analysis (ensemble forecasting).
- The nonlinear dynamics of numerical models makes perturbations grow exponentially in time and localize in space.
- There are different methods to perturb the initial conditions and the spatio-temporal growth of the perturbations differs from one technique to the other.
- The characterization of perturbations growth contributes to determine the predictability barrier of the system.

Experiment:

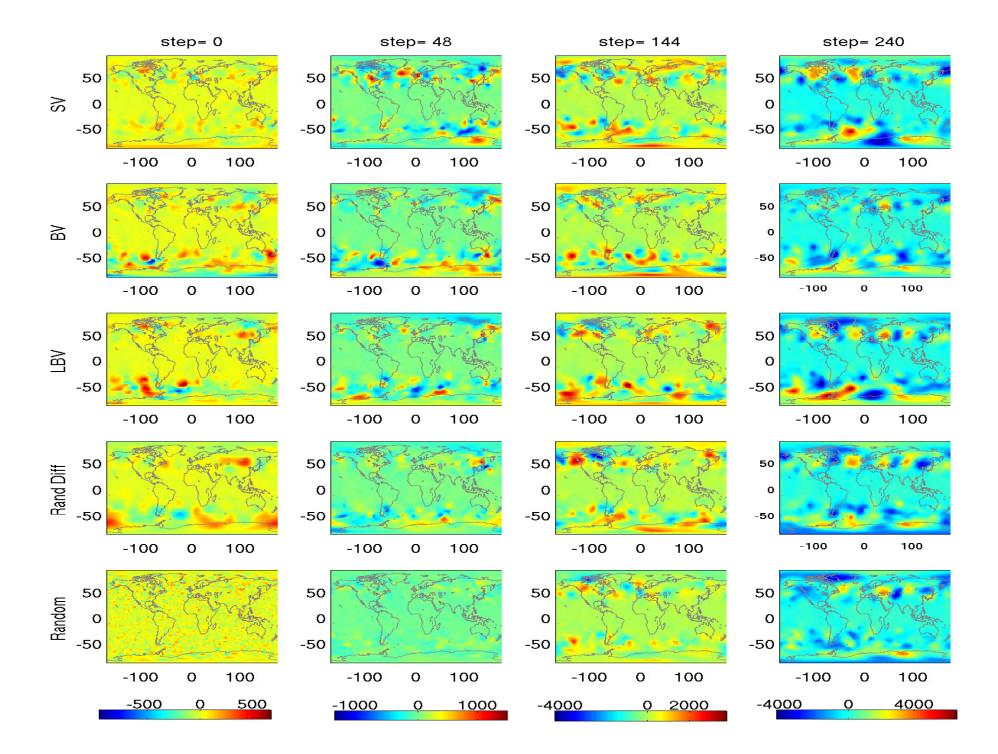
Spatio-temporal comparison of initial perturbation techniques:

- The operational singular vector system at ECMWF (SV), Lorenz 1965, Palmer et al (1993)
- Breeding vectors (BV) Toth and Kalnay (1993,1997)
- Breeding vectors rescaled by a geometrical mean (LBV) Primo et al (2005)
- Random field perturbation (RF) Magnusson et al (2008).
- Random perturbation technique

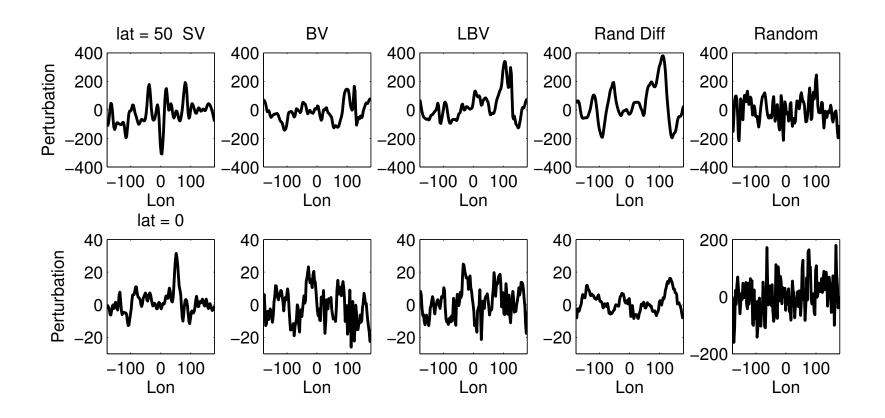
Experiment

- Data: Generated by Linus Magnusson (Stockholm University)
- Model: ECMWF Ensemble Prediction System.
- Period: 20050101-20050115
- Number of Members: 20
- Steps: from 0h to 240h by 12 h.
- Parameter: Geopotential
- Level: 500HPa
- Time: 00:00:00
- Different areas:

```
North Hemisphere (90°N, 30°S, -180°W, 180°E),
South Hemisphere (-30°N, -90°S, -180°W, 180°E),
Europe (75°N, 35°S, -12.5°W, 42.5°E)
Tropics (-30°N, 30°S, -180°W, 180°E).
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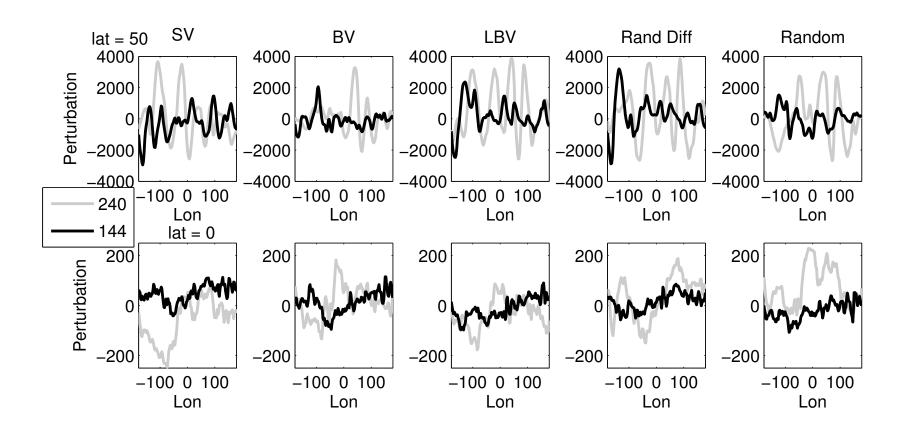
Initial perturbations



Perturbations are dominated by the extratropics, but the tropics also show spatial structure

Perturbations evolution

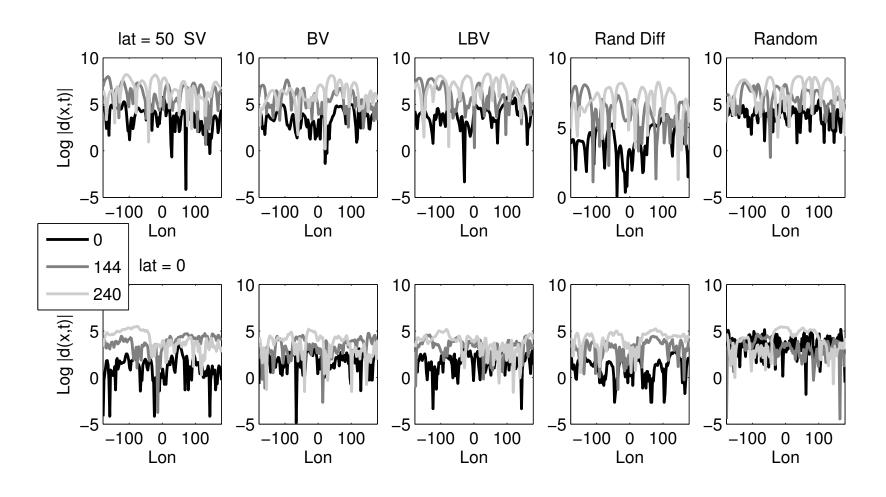
Perturbations after 6 (black) and 10 (grey) days



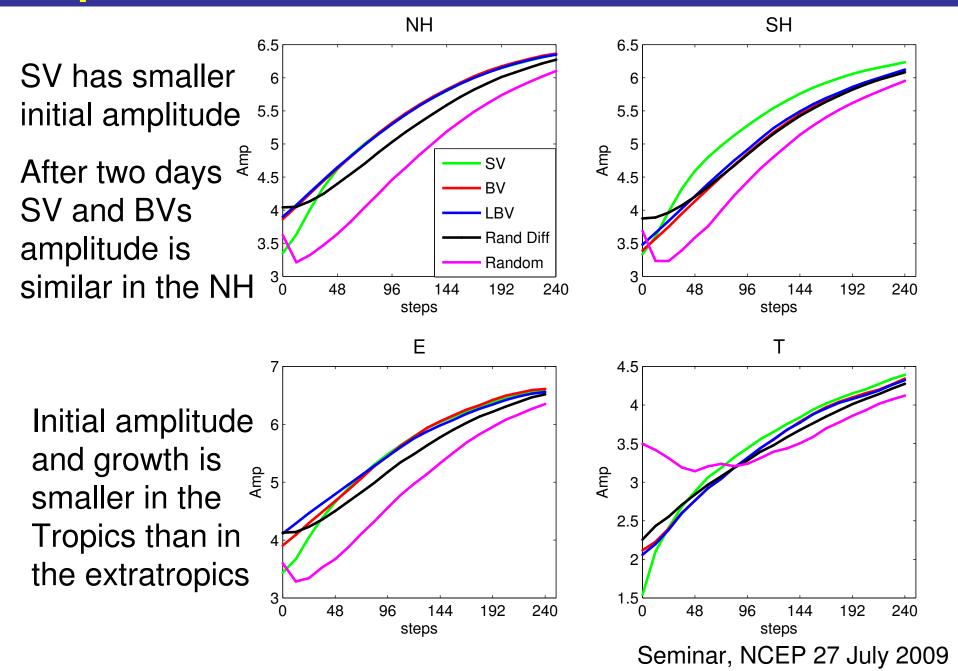
Spatial structures become bigger and tend to a similar pattern

Logarithmic perturbations evolution

 $h(x,t) = \ln |\delta u(x,t)|, x = 1,...,L$ Hopf - Cole transformation

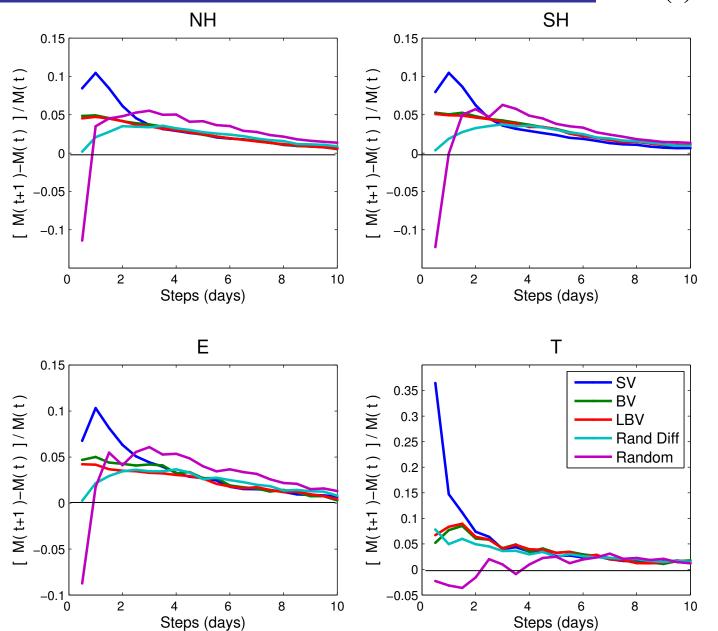


Amplitude Growth



Amplitude Growth rate:

$$Rate(t) = \frac{M(t+1) - M(t)}{M(t)}$$



SV has fastest initial growth, but after around three days growth rate converges

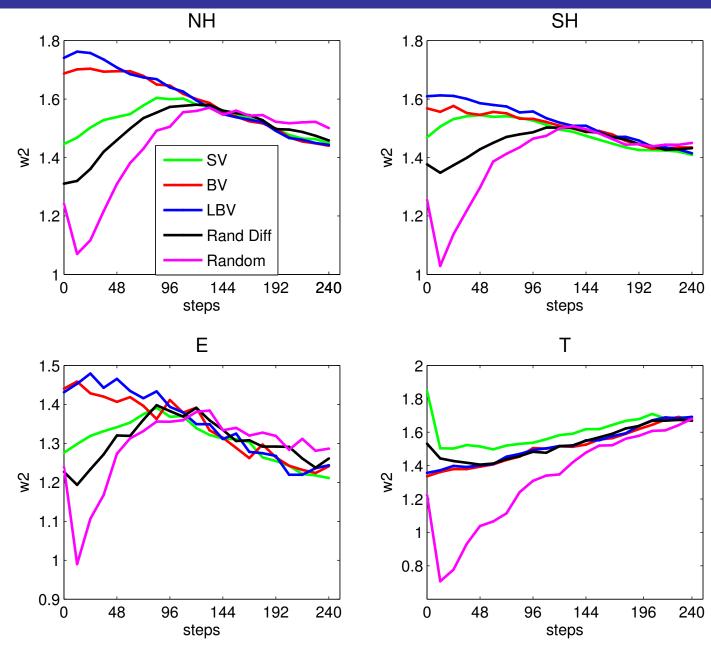
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Variance growth

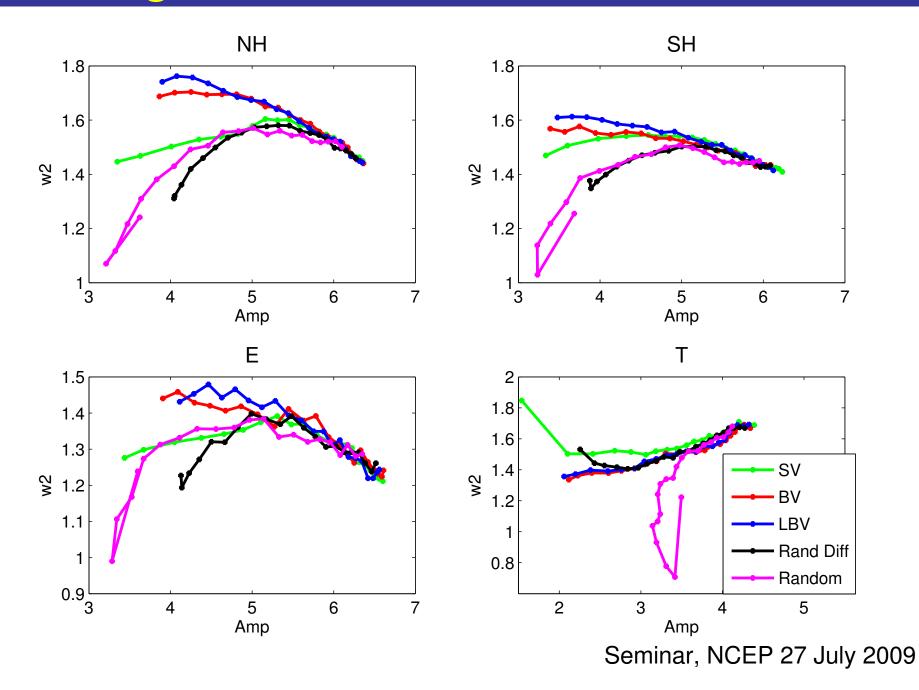
BVs provide perturbations with the highest correlation in the extratropics

Random perturbations start with lower correlation, but after around five days all techniques converge

Correlation in the Tropics seems not to be saturated after 10 days

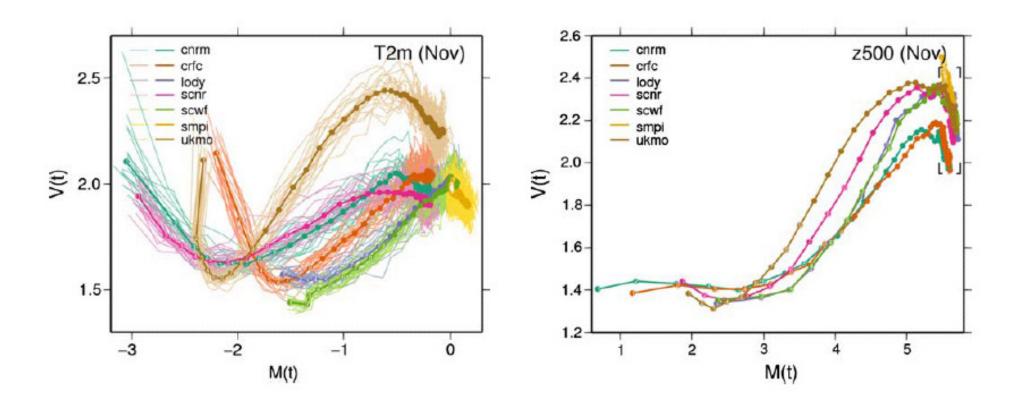


MVL Diagram



Seasonal forecasts: DEMETER

 The DEMETER project is a multi-model seasonal ensemble with hindcasts for seven models covering a common 22-year period (1980-2001).



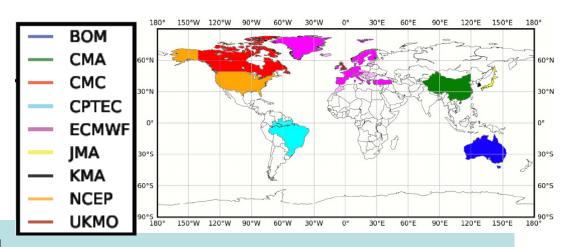
Fernández et al. (2009, figs 2, 3)

TIGGE project provides an archive of medium-range ensemble forecasts from a range of different operational models in a consistent format.

Centre	Location	Horiz. res.	Levels	Members	Perturbations
BOM	Australia	T119	19	1 + 32	SV (initial)
CMA	China	T213	31	1 + 14	BV
CMC	Canada	0.9°	58	20	EnKF (since 2005)
CPTEC	Brazil	T126	28	1 + 14	EOF
ECMWF	Europe	T399	62	1 + 50	SV (initial+final) + stoch.
JMA	Japan	T319	60	1 + 50	SV (since Nov 2007)
KMA	South Korea	T213	40	1 + 16	BV
NCEP	USA	T126	28	1 + 20	ETR + stoch.
UKMO	UK	$1.25^{\circ} \times 0.833^{\circ}$	38	1 + 23	ETKF + IAU + stoch.

BV Bred Vectors
EOF Empirical Orthogonal Functions
EnKF Ensemble Kalman Filter
ETKF Ensemble Transform Kalman Filter
ETR Ensemble Transform with Rescaling
IAU Incremental Analysis Update
SV Singular Vectors (initial- and/or final-time)
stoch. Stochastic parameterisation

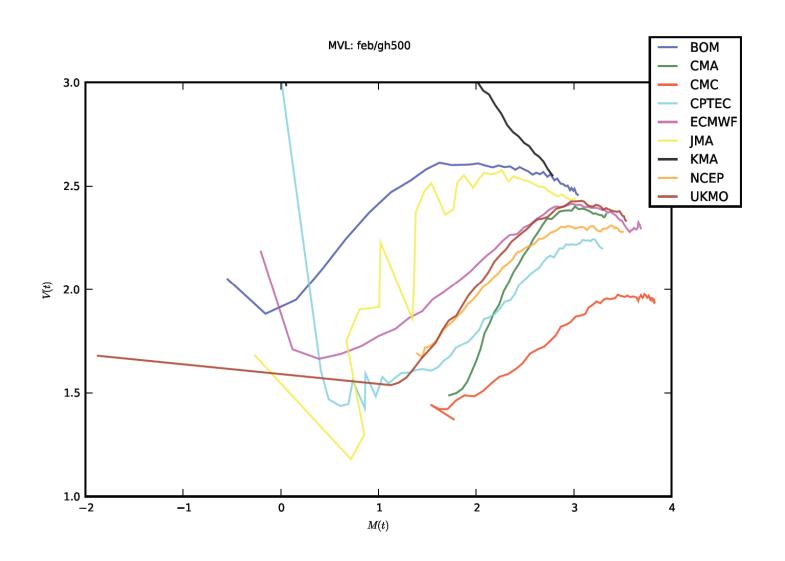
The regions of the globe over which perturbations are optimised differ from model to model

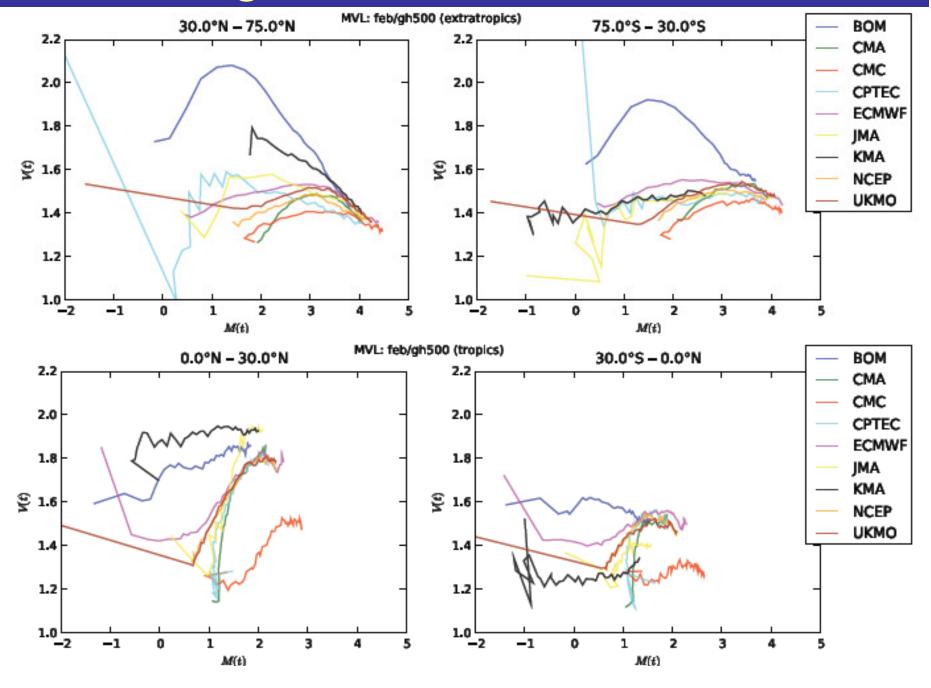


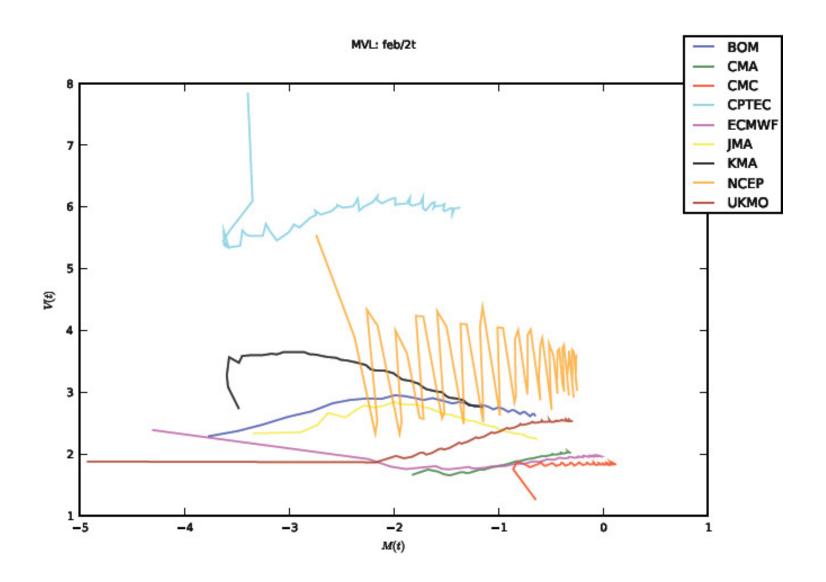
MSc Student: Zak Kipling

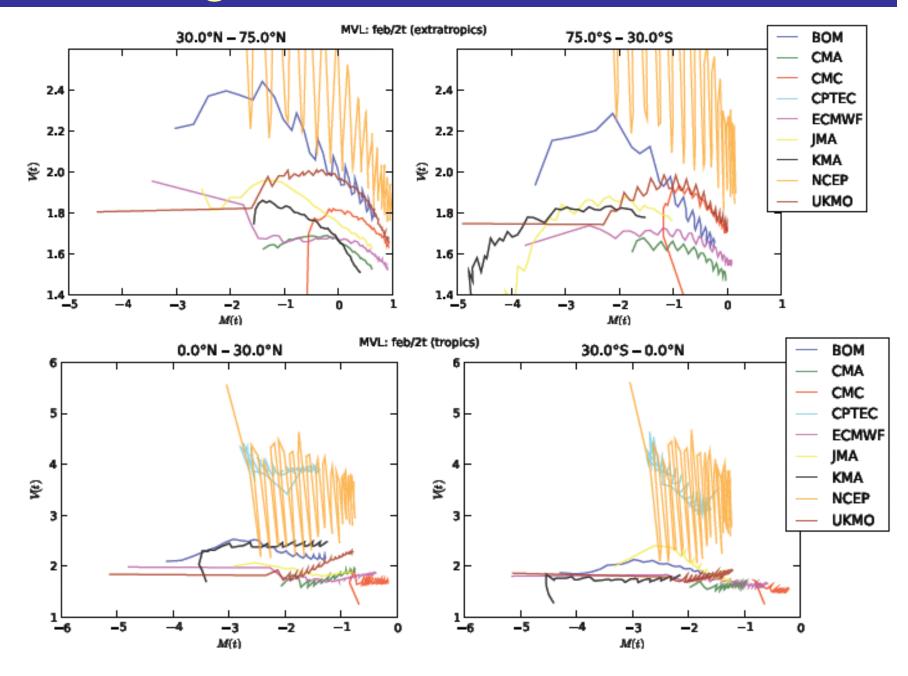
Project: Error growth in medium-range forecasting models

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Summary and conclusions

- Perturbations have an exponential growth, get spatially correlated and localized.
- A logarithmic transformation allows us to link perturbations and rough interfaces growth, providing scaling properties already obtained in this field.
- Different initial perturbation techniques have been compared using the MVL diagram.
- This diagram is shown to be an useful tool to analyze the spatiotemporal growth of perturbations, distinguising the linear and non linear regime and the correlation and amplitude growth.
- When comparing different initial perturbation techniques applied to the ECMWF EPS, SV and BV have similar amplitude after 2 days in the North Hemisphere, and similar spatial correlation after 4-5 days.
- Perturbations in the Tropics grow more slowly and their spatial correlation seems not to be saturated after 10 days for any of the techniques.

References

- Pikovsky, A. S. and Kurths, J. (1994): "Roughening interfaces in the dynamics of perturbations of spatiotemporal chaos". Physical Review E 49(1), 898-901.
- Pikovsky, A. and Politi, A. (1998): "Dynamic localization of Lyapunov vectors in spacetime chaos". Nonlinearity 11(4), 1049-1062.
- López, J.M., Primo, C., Rodríguez, M.A., and Szendro, I. G. (2004): "Scaling properties of growing non infinitesimal perturbations in space-time chaos". Phys. Rev. E, 70, 056224.
- Primo, C., Szendro, I. G., Rodríguez, M. A. and Gutierrez, J. M. (2007): "Error growth patterns in systems with spatial chaos: From coupled map lattices to global weather models". Physical Review Letters 98(10).
- Gutierrez, J. M., Primo, C., Rodriguez, M. A. and Fernandez, J. (2008): "Spatiotemporal characterization of ensemble prediction systems the mean-variance of logarithms (MVL) diagram", Nonlinear Processes in Geophysics 15(1), 109-114.
- Palmer, T. N. (1999): "Predicting uncertainty in forecasts of weather and climate". Technical Memorandum 294, ECMWF.
- Lorenz E. N. (1965): "A study of predictability of a 28-variable atmospheric model", Tellus, 67, 321-333.
- Toth, Z. and Kalnay E. (1993): "Ensemble forecasting at NCEP: The generation of perturbation". Bull. Am. Meteor. Soc, 74, 2317-2330.
- Toth,Z. and Kalnay E.(1997): "Ensemble Forecasting at NCEP and the breeding method".
 Mon. Weather Rev. 125, 3297-3319.
- Primo C., Rodríguez M.A., Gutiérrez J.M. (2008): "Logarithmic Bred Vectors. A New Ensemble Method with Adjustable Spread", Journal of Geophysical Research, 113, D05116
- Magnusson, L., Nycander, J. and Källen, E.(2008): "Flow-dependent versus flow-independent initial perturbations for ensemble prediction", Tellus, 61A, 194-209.
- Fernandez, J., Primo, C., Cofiño, A., Gutierrez, J. and Rodríguez, M. (2009): "MVL spatiotemporal analysis for model intercomparison in EPS: application to the DEMETER multi-model ensemble". Climate Dynamics 33(2), 233-243.

Any questions?